

# Complete Observables for Canonical General Relativity

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## Overview

1. One Constraint
2. Arbitrary Many Constraints
3. General Relativity: infinitely many constraints
4. Reducing the Number of Constraints:  $\infty \rightarrow$  one
5. Outlook

# One Constraint: FRW–Cosmology

Method: C.Rovelli (1990/2002)

Example: FRW–cosmology with massless scalar field

phase space:  $x = (a, P_a, \phi, P_\phi)$       constraint:  $C = \frac{1}{2} \left( -\frac{P_a^2}{a} + \frac{P_\phi^2}{a^3} \right)$

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Dirac Observable for arbitrary  $\tau$

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⇒ To compute this introduce a new basis of constraints  $\check{C}_i$

$$\check{C}_i = C_k (A^{-1})_{ki} \quad \text{with} \quad A_{lm} = \{T_l, C_m\}$$

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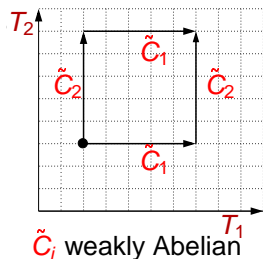
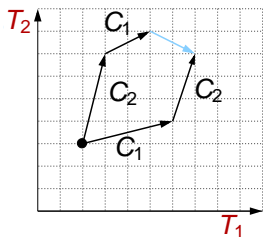
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# Power Series

$$\mathbf{F}_{[f; T_i]}^{T_i}(x) = \sum_{k_j=0}^{\infty} \frac{1}{|k|!} \tilde{C}_1^{k_1} \cdots \tilde{C}_N^{k_N} [f](x) (\tau_1 - T_1(x))^{k_1} \cdots (\tau_N - T_N(x))^{k_N}$$



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acts on  $f$ :  $\tilde{C}_i[g] := \{g, \tilde{C}_i\}$

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'distance' of  $x$  to  $T_i = \tau_i$

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:=  $f(x')$  with

$x'$  gauge equivalent to  $x$  and

$$T_i(x') = \tau_i \quad \forall \quad i$$

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## Some Results

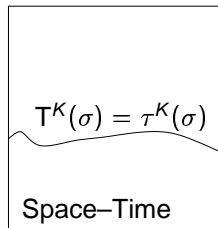
- complete set of Dirac observables
- Poisson brackets between complete observables
- (  $\mathbf{F}_{[\cdot; T_i]}^{T_i}$  is a symplecto-morphism onto space of Dirac observables)
- physical Hamiltonians generate evolution wrt clock variables  
(Kuchař 1972, Rovelli 1990, Thiemann 2004, B.D 2005)
- partial observables invariant under subset of the constraints:  
“non-perfect clocks”

- infinitely many constraints:  $C_a(\sigma)$ ,  $a = 1, 2, 3$ ;  $C_\perp(\sigma)$   
 $\sigma \in \Sigma$  (spatial hypersurface)

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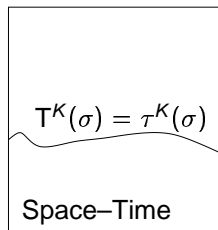
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- $T^K(\sigma) = \tau^K(\sigma)$  fix embedding of  $\Sigma$  into space-time
- $\mathbf{F}_{\tau^K(\sigma)}^{T^K(\sigma)}[f; T^K(\sigma)]$  = value of  $f$  on this embedding
- example:  $f = \phi(\sigma^*)$

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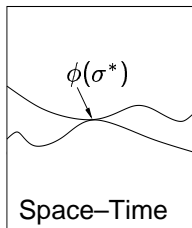


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- infinitely many parameters needed to fix embedding
- but: choose  $f$  such that it does not depend on all aspects of the embedding

↪  $f$  invariant under a subset of the constraint:  $\rightarrow$  space-time scalar

# Reducing the Number of Constraints

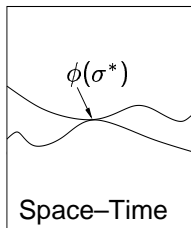


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 $\{C[N, N^a], \phi(\sigma^*)\} = 0$  for  $N(\sigma^*) = 0 = N^a(\sigma^*)$

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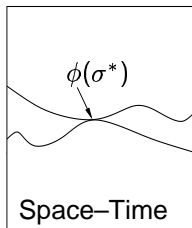


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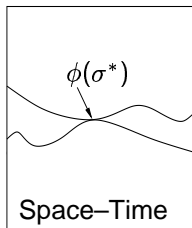
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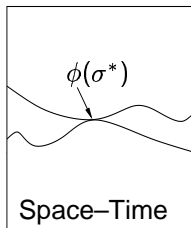
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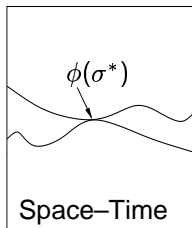
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**Need only four constraints for calculation of complete observable**

$$F_{[\phi(\sigma^*); T^K(\sigma)]}^{\tau^K(\sigma^*)} = \sum_{k_K=0}^{\infty} \frac{1}{|k|!} \tilde{C}_0^{k_0}[1] \cdots \tilde{C}_3^{k_3}[1][\phi(\sigma^*)] (\tau_0 - T^0)^{k_0} \cdots (\tau_3 - T_3)^{k_3}$$

depends on only four parameters  $\tau^K(\sigma^*)$

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  - ★ are canonically conjugated
  - ★ physical Hamiltonian:  $\{\mathbf{F}_{\gamma^{KL}}^{\tau^K}, H_M^{\tau^K(\sigma)}\} = \mathbf{F}_{\partial_M \gamma^{KL}}^{\tau^K} = \frac{\partial}{\partial \tau_M} \mathbf{F}_{\gamma^{KL}}^{\tau^K}$

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**Need only one constraint for calculation of complete observable!**

choose another scalar  $T^0(\tau^C)$ :  $\tilde{C}(\tau^C) = [\{T^0(\tau^C), C_{\perp}[1]\}]^{-1} C_{\perp}(\tau^C)$

$$\mathbf{F}^{\tau^C}_{\gamma^{KL}(\tau^C)} = \sum_k \frac{1}{k!} \tilde{C}[1]^k [\gamma^{KL}(\tau^C)] (\tau^0 - T^0(\tau^C))^k$$

# Conclusions and Outlook

- partial and complete observables: method to systematically investigate observables and their algebra
- can express observables of the covariant formalism in the canonical formalism: compare quantizations
- reduction to one constraint  $\rightsquigarrow$  approximation scheme
- diffeomorphism invariant Abelian Hamiltonian constraints without square roots

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  - can express observables of the covariant formalism in the canonical formalism: compare quantizations
  - reduction to one constraint  $\rightsquigarrow$  approximation scheme
  - diffeomorphism invariant Abelian Hamiltonian constraints without square roots
- 
- ★ good clock variables?  $\rightsquigarrow$  physical Hamiltonian (positive,  $\tau$ -independent)
  - ★ how to deal with bad clock variables?
  - ★ quantization: different clock variables (with T. Thiemann)

# Appendix

$$\begin{aligned} C_{\perp}(\tau^C) = & \frac{1}{\kappa} g^{-\frac{1}{2}} (p_{AB} p^{AB} - \frac{1}{2} p^2)(\tau^C) - \frac{1}{\kappa} g^{\frac{1}{2}} R(\tau^C) + \\ & \frac{1}{2\alpha_0} g^{-\frac{1}{2}} (\Pi_0)^2(\tau^C) + \frac{1}{2\alpha_0} g^{\frac{1}{2}} g^{AB} T_{,A}^0 T_{,B}^0(\tau^C) + \\ & \frac{1}{2} g^{-\frac{1}{2}}(Y) \sum_{A=1}^3 \frac{1}{\alpha_A} (\Pi_A)^2(\tau^C) + \frac{1}{2} g^{\frac{1}{2}}(Y) \sum_{A=1}^3 \frac{1}{\alpha_A} g^{AA}(\tau^C) + \\ & g^{\frac{1}{2}} \sum_{K=0}^3 \frac{1}{\alpha_K} V_{(K)}(T^0(\tau^C), \tau^C) \end{aligned}$$