

QUANTUM  
BLACK HOLES:  
DYNAMICS.

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S to be posted.

LOOPS 05

THE PROBLEM:

FIND FULL QT OF THE GRAVITY-SCALAR  
FIELD THEORY IN SPHERICAL SYMMETRY.  
(ASYPT. FLAT SECTOR)

NON-TRIVIAL 1+1 DIM. FIELD THEORY

WHY?

- \* WHAT DOES QG COLLAPSE LOOK LIKE?  
CLASSICAL COLLAPSE WELL STUDIED  
NUMERICALLY (GOLDWIRTH-PIRAN, CHOPTUIK ...)
- \* WHAT IS A "QUANTUM BLACK HOLE"?  
SEEK THE RESULT OF A LONG TIME  
QUANTUM COLLAPSE EVOLUTION.
- \* DOES HAWKING RADIATION SHOW  
UP IN A SUITABLE APPROXIMATION?  
RESOLUTION OF TRANSPALANKIAN MODE  
PROBLEM?

②

THIS PROBLEM WAS POSED BY UNRUH (1976) TO UNDERSTAND HAWKING RADIATION.

→ FULL GAUGE FIXING TO GET A REDUCED THEORY FOR THE SCALAR FIELD:

$$H[\alpha, P_\alpha] = \int_0^\infty dr f(\alpha, P_\alpha) e^{\int_0^r dr' g(\alpha, P_\alpha)}$$

" I PRESENT IT HERE IN THE HOPE THAT SOMEONE ELSE MAY BE ABLE TO DO SOMETHING WITH IT "

\* USELESS HAMILTONIAN IS THE RESULT OF A RADIAL COORDINATE GAUGE FIXING.

OUR STRATEGY:

- \* FIX A TIME GAUGE TO BE ABLE TO ADDRESS AT LEAST SOME DYNAMICAL QUESTIONS UNAMBIGUOUSLY
- \* AVOID A RADIAL COORD. GAUGE FIXING TO GET A MORE TRACTABLE HAMILTONIAN.

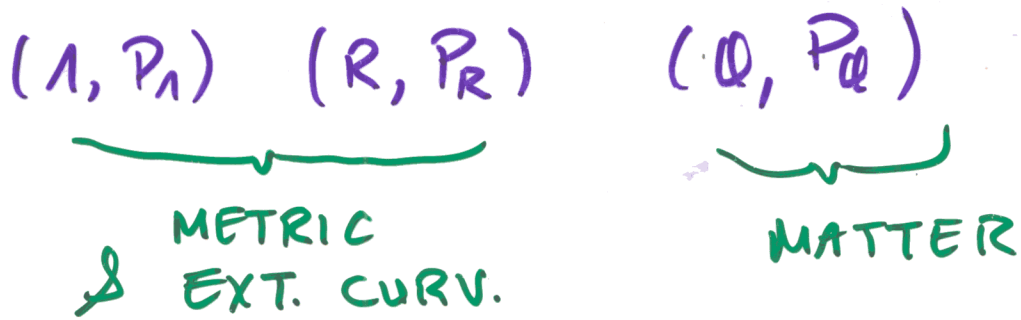
1+1 THEORY FOR METRICS

$$ds^2 = -(N^2 - (N^r)^2) dt^2 + 2N^r dr dt + r^2 dr^2 + R^2 d\Omega^2$$

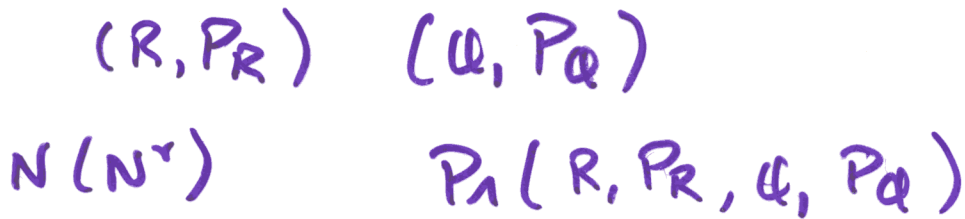
WITH

$$T_{ab} = \partial_a \mathcal{Q} \partial_b \mathcal{Q} - \frac{1}{2} g_{ab} (\partial \mathcal{Q})^2$$

1+1 ADM VARIABLES



\* TIME GAUGE FIXING →



\* REDUCED ACTION:

$$S_R = \int dr dt \left[ P_\phi \dot{\phi} + P_R \dot{R} - N^r (P_1' + P_R R' + P_\phi \phi') \right] + \int dt (P_1 N^r)$$

CLASSICAL EVOLUTION :

EQNS  $\dot{\Phi} = \{ \Phi, \mathcal{H}(N^r) \}$  etc.

CONTAIN TWO PIECES :

$$\dot{\Phi} = \underbrace{\text{EVOLUTION IN TIME}} + \int_{N^r} \mathcal{Q} .$$

GENERATED BY

$$H = \int dr (N^r)' P_{\Lambda} \neq 0$$

WE WOULD LIKE  $H \rightarrow \hat{H} \dots$

BUT

$$P_{\Lambda} = P_R R + \sqrt{(P_R R)^2 - 16 R^2 (H_{\mathcal{Q}} - {}^{(3)}R)} \quad !$$

$$\left[ \begin{array}{l} H_{\mathcal{Q}} = \frac{P_{\mathcal{Q}}^2}{2R^2} + \frac{R^2}{2} \mathcal{Q}'^2 \\ {}^{(3)}R(R, R', R'') = \text{Ricci scalar of } g_{ab} \end{array} \right]$$

# A CONSTRUCTION OF $\hat{H}$ BY LINEARISATION OF $\sqrt{\quad}$ :

$$* \quad \hat{\sqrt{\quad}} = v_1 \hat{A} + v_2 \hat{B} + v_3 \hat{C} + v_4 \hat{D}$$

(A, ... D) ARE THE 4 PIECES IN ROOT)

$v_i$  ARE VECTORS IN A SUBSPACE OF  $4 \times 4$  MATRICES SATISFYING

$$v_i \cdot v_j = \eta_{ij} I = \text{diag}(1, -1, -1, 1) I$$

\*  $v_i$  ARE REAL MATRICES.

\* QUANTUM STATES ARE REAL SPINORS:

EG.  $\begin{pmatrix} a \\ b \\ c \\ d \end{pmatrix} | \underbrace{a_1 \dots a_n}_{\text{R excitations}} ; \underbrace{b_1 \dots b_n}_{\text{excitations}} \rangle$

(FROM OLIVER'S TALK)

WITH  $\hat{H}$  DEFINED WE HAVE

$$i \frac{\partial}{\partial t} |\psi\rangle = \hat{H}(N^r) |\psi\rangle$$

AND

$$\hat{C}(N^r) |\psi\rangle = 0$$

\* WE CAN TRACK EVOLUTION (IN GAUGE FIXED TIME) WITHOUT SOLVING THE SECOND EQN, FOR GIVEN  $N^r$  (WHICH SATISFIES ASYPT. CONDITIONS).

of YM theory:

$$i \frac{\partial}{\partial t} |\psi\rangle = [\hat{H}_{YM}(E, A) + \lambda \cdot \hat{G}] |\psi\rangle$$

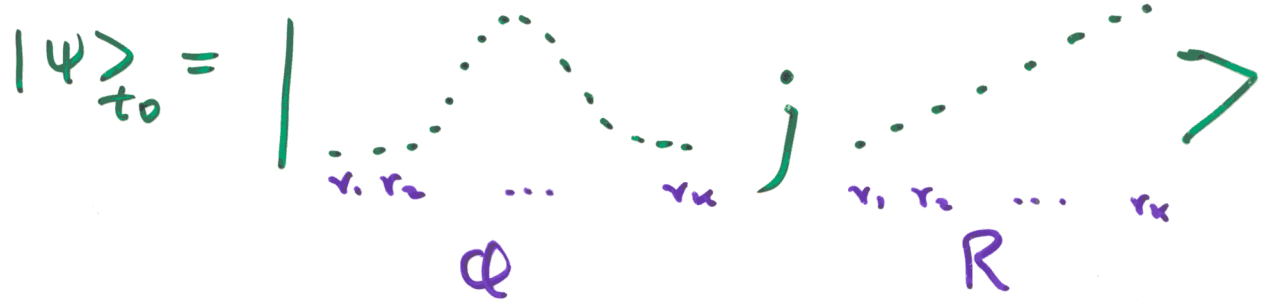
$$\hat{G} |\psi\rangle = 0$$

--- EVOL OF  $|\psi\rangle$  POSSIBLE WITHOUT SOLVING  $\hat{G} |\psi\rangle = 0$  FOR PRESCRIBED  $\lambda$ .



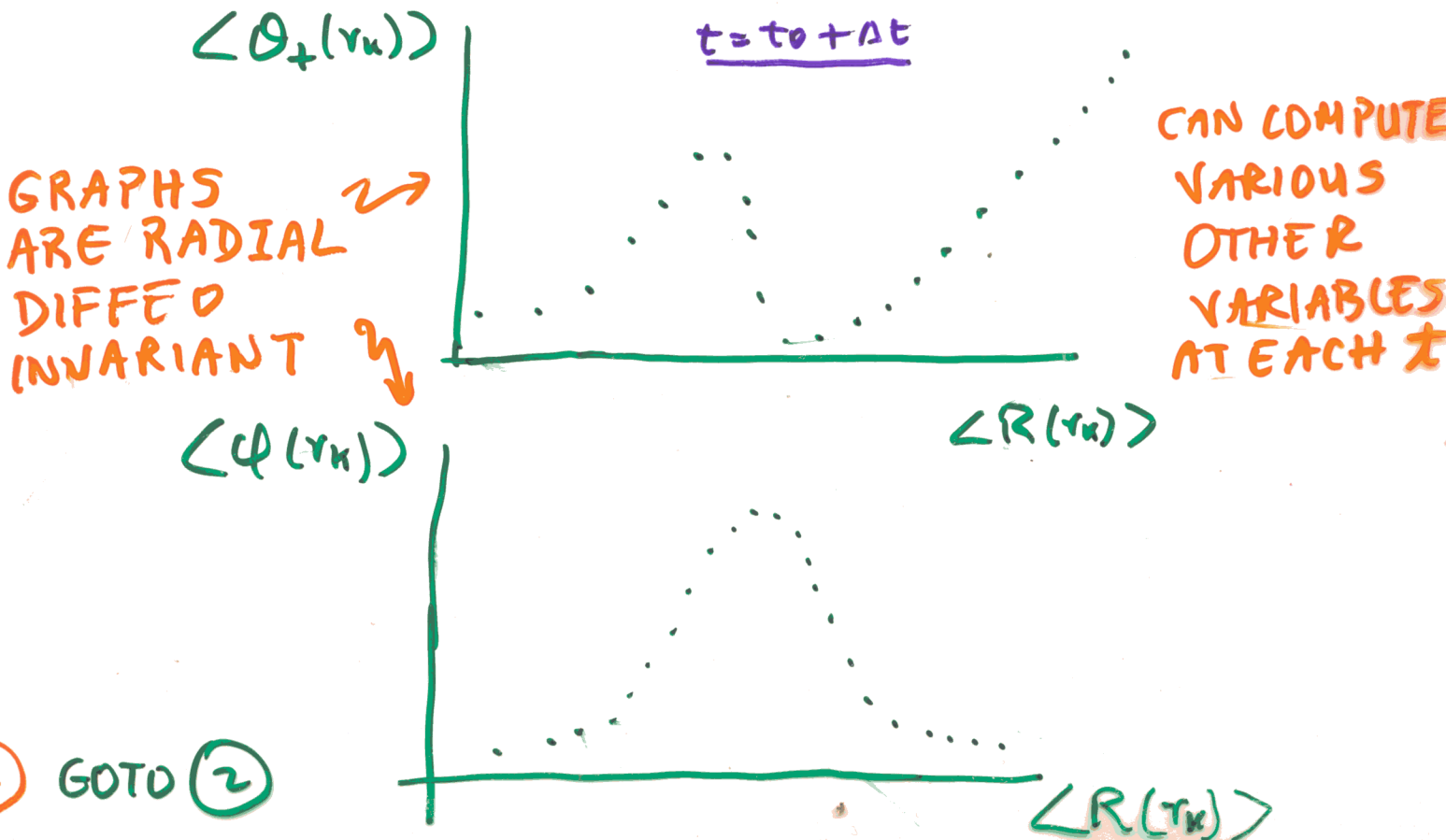
ALGORITHM FOR COLLAPSE EVOL :

① PICK INITIAL STATE



②  $|\psi\rangle_{t_0+\Delta t} = (I + i\Delta t \hat{H}) |\psi\rangle_{t_0}$

③ COMPUTE :



SUMMARY/COMMENTS:

- \* AN EXPLICIT SETUP FOR STUDYING GRAV. COLLAPSE IN QG IN A FIELD THEORY SETTING
- \* BY HAND CALCULATION FOR GRAPH OF 3 POINTS: SEE MOVEMENT OF  $\langle Q_+(Y_u) \rangle$ ,  $\langle Q(Y_u) \rangle$  vs  $\langle R(Y_u) \rangle$  FROM  $t_0 \rightarrow t_0 + \Delta t$ .
- \* REALISTIC EVOLUTION REQUIRES NUMERICAL WORK.  
EASIER THAN CORRESPONDING CLASSICAL COLLAPSE IMPLEMENTATION.  
IN PROGRESS (... LOOPS OF)



- \* OTHER GAUGES
- \* PARTIAL OBSERVABLES
- \* GAUGE FIXED SPIN NETWORK EVOL
- \* BOSONISATION