

Sanford Sharpie.

Order + Number → Geometry

LumiColor Orange

Lumi Green

Sanford Green

Stabilo Black

## 1. Order + Number $\Rightarrow$ Geometry

Discreteness ADDS information too (as well as removing degrees of freedom)

Volume = number

## 2. Discreteness and Lorentz invariance can live together

In fact  $N=V$  requires Lor invar

$\Rightarrow$  kinematic randomness

$\Rightarrow$  UV-IR mixing

$\Rightarrow$  locality must be given up (tho not causality)

Recovering approximate locality  $\Rightarrow$  *intermediate length scale* enters on which nonlocality survives (cf non-commutative geometry)

[By the way, nonlocality arguably does *not* cure the perturbative infinities of qft, it worsens them! The *cutoff* cures them (but they're related)]

### 3. The “Problem of Time” is soluble in a discrete, past-finite cosmos

A *background-free* generally covariant dynamics exists.

Its “observables” (label-invariant predicates) are not only formally defined, but *physically accessible* (in terms of “stem predicates”)

(This dynamics is classical, but covariance and background-independence aren’t issues for the quantal generalization)

### 4. Covariance and becoming can co-exist: *dynamics as growth*

5. The Cosmological Constant might be a (nonlocal & quantal) residue of the underlying discreteness.

(i)  $\Lambda$  was predicted as a fluctuation effect ( $1/\sqrt{N}$ )

(if so  $d = 4$  is special, the “critical dimension”)

The supernovas are illuminating the underlying discreteness!

(ii) “Why continuum?” = “why is  $\Lambda \approx 0?$ ”

(nonlocality might be key *together with* quantal interference)

(say why!)

6. The large numbers of cosmology might be understood without inflation as reflecting the large age of cosmos: *its many cycles*

→ a new kind of renormalization induced by “bounce”

7. Causality + Covariance is restrictive → CSG

(causality in derivation of CSG is replacing locality in derivation of GR)

we hope they will also lead us to QSG (if not then what? entropy bound?  $\square\square\sigma?$ )

## 8. Quantum Gravity needs Generalized Quantum Mechanics

It needs a histories formulation that can do without unitarity

QM as a generalized form of measure theory

We shouldn't expect a unitary theory if we freeze the causet (resp manifold) This is so *even if* full theory were unitary, but:

We shouldn't expect unitarity in full QG either

## 9. BH entropy might be understood as “number of horizon molecules”

(cf LQG)

We also see hint of “holographic entropy bound”

10. Discreteness can have phenomenological consequences even while respecting Lorentz-invariance

⇒ diffusion in velocity space

⇒ deviations from  $\square\phi = 0$   
(maybe even too big)

⇒  $\Lambda$  fluctuations  
(incidentally, excludes “large extra dimensions”)

# DISCRETE GRAVITY

Why discreteness?

The infinities

- QFT  $\Rightarrow \frac{1}{2} h\nu$  per mode  $\Rightarrow S \approx \infty$  ( $\Lambda = \infty$ )  
(black body)
- GR  $\Rightarrow$  Singularities where Curvature  $\rightarrow \infty$   
 $\Rightarrow$  structure on infinitesimal scales
- std. model (prob) lacks continuous limit  
 $\Rightarrow \nexists$  continuum?
- QFT + GR is not perturbatively renormalizable  
(gravity does not cure  $\infty$ 's of QFT  
if continuum retained)  
 $\Rightarrow$  something new at  $\ell_{\text{Pl}}$ , but what?  
(&  $\nexists$  SRDV must fail even earlier)
- GR + QFT  $\Rightarrow \infty$  BH "entanglement entropy"  
(cutoff cures it)
- Continuum  $\Rightarrow \infty$  # of horizon shapes  
which again  $\Rightarrow S_{\text{BH}} = \infty$   
(we don't expect pure state for these vbls.)

Problem is  $\infty$ 'ly small distances

The alternative continuum-discontinuum seems to me to be a real alternative, i.e. there is no compromise... In a [discontinuum] theory space and time cannot occur... It will be especially difficult to derive something like a spatio-temporal **quasi-order** from such a schema... But I hold it entirely possible that the development will lead there...

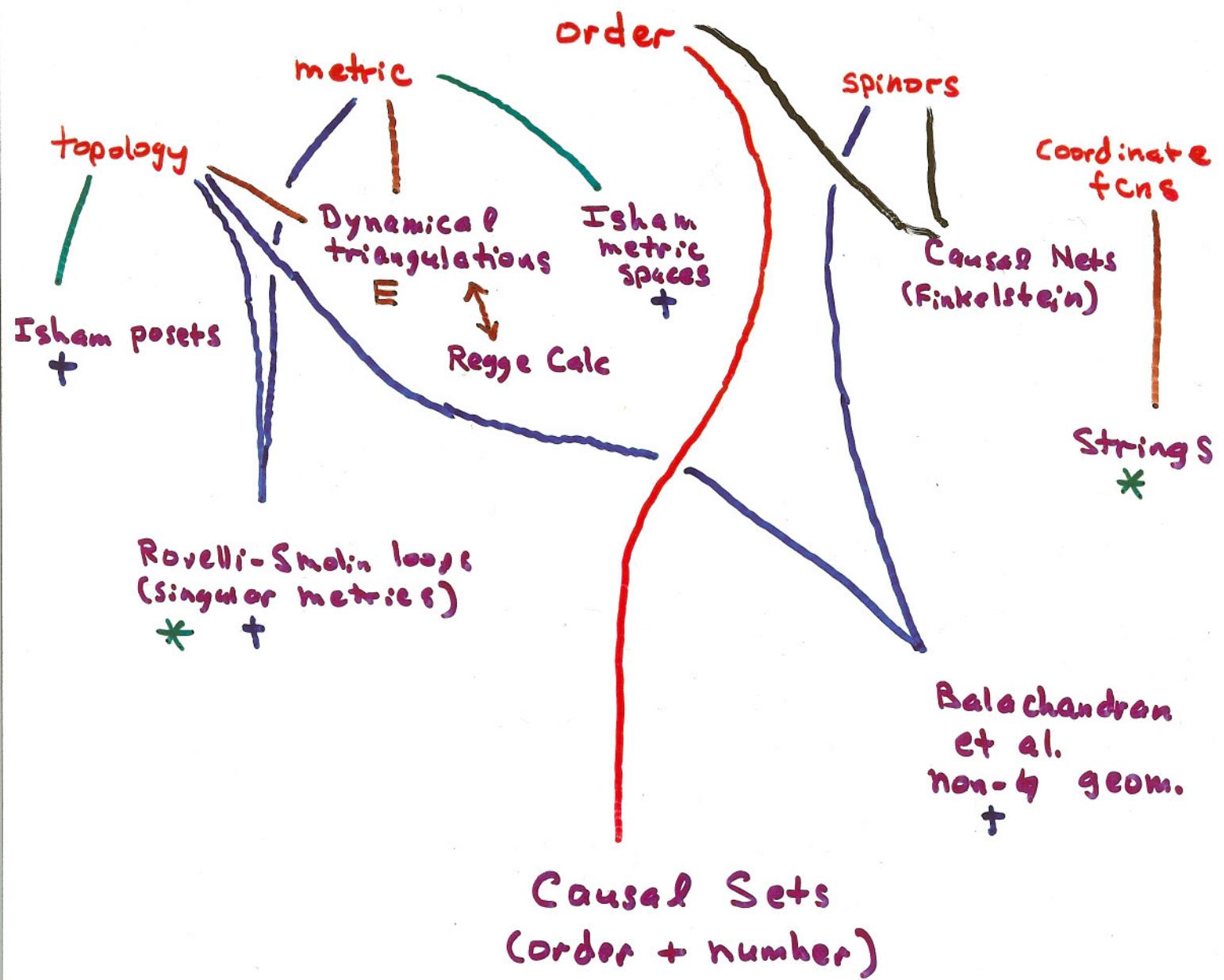
1954

Determinate parts of a manifold ... are called quanta. Their quantitative comparison occurs, for discrete [manifolds] through counting, for continuous ones through measurement.

1854

What is the underlying discrete structure?

5 bridges back to the continuum



\* = manifold retained

t = Only space recovered, and  $t \in \mathbb{R}$

E = Euclidean Signature

The causal order is arguably  
the most fundamental  
Spacetime structure,  
and the simplest.

From it we can get all  
the others (with discreteness!)

# Introduction to Causets

A kind of "(modern) family tree"  
(also early?)



example

$N \leftrightarrow \text{Space} \underline{\text{time}} \text{ Volume}$

microscopic  
order  $\leftrightarrow$  macroscopic  
causal order

(light cones)

Together:

Number + Order  $\Rightarrow$  Geometry

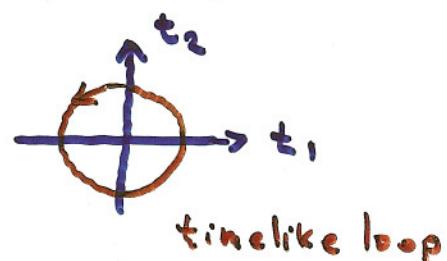
discreteness is essential to get volume element

Unifies: topology, diff. str., metric  
in terms of order

(can it yield "matter" too?)

Answers: why length?

explains  $(- +++)$   
 $\underbrace{+ +}_{D-1}$



Also:

Coarse-graining allows description  
of (eg) scale-dependent topology  
(Kaluza-Klein, "foam")

$\Rightarrow$  no closed timelike curves

We can prove this:

- (i) If  $C$ 's partial stems do not characterize it then it must contain  $\infty$  copies of some stem  $S$ .
- (ii) If some  $S$  occurs infinitely often in  $C$  then some level of  $C$  is  $\infty$
- (iii)  $C$  almost surely has no  $\infty$  levels if  $t_2 > 0$ .  
(the exceptions: antichain, tree are harmless)

This gives physical meaning to our formal  $\sigma$ -algebra  $\mathcal{R}$ .

A related issue to consider  
"this partial stem" vs "some partial stem"

# Fluctuating Cosmological "Constant"

$$(\delta \Lambda \sim V^{-1/2}) \text{ (Heuristic)}$$

In seq. growth dynamics,  $N$  plays formally role of time :  $N \leftrightarrow T$

But we normally don't integrate over  $T$  ; in  $\Sigma$ /histories, we hold it fixed.  
 $\Rightarrow$  expect to fix  $N$  in " $\Sigma$ /causets"

$N \leftrightarrow V \Rightarrow$  would fix  $V$  in Continuum approx.  
 (unimodular)

But this  $\Rightarrow \Lambda$  free:

$$\delta \left( \int \left( \frac{1}{2\pi} R - \Lambda_0 \right) dV - \lambda V \right) = 0$$

$\Rightarrow$  only  $\Delta = \Lambda_0 + \lambda$  matters

Quantum:  $\Delta, V$  conjugate  $\Rightarrow \delta \Lambda \delta V \sim k$

But Poisson  $\Rightarrow N = V \pm \sqrt{V} \Rightarrow V \approx N \pm \sqrt{N}$

i.e.  $\delta V \sim \sqrt{V}$

$$(*) \Rightarrow \delta \Delta \sim \frac{k}{\delta V} \sim \frac{1}{\sqrt{V}}$$

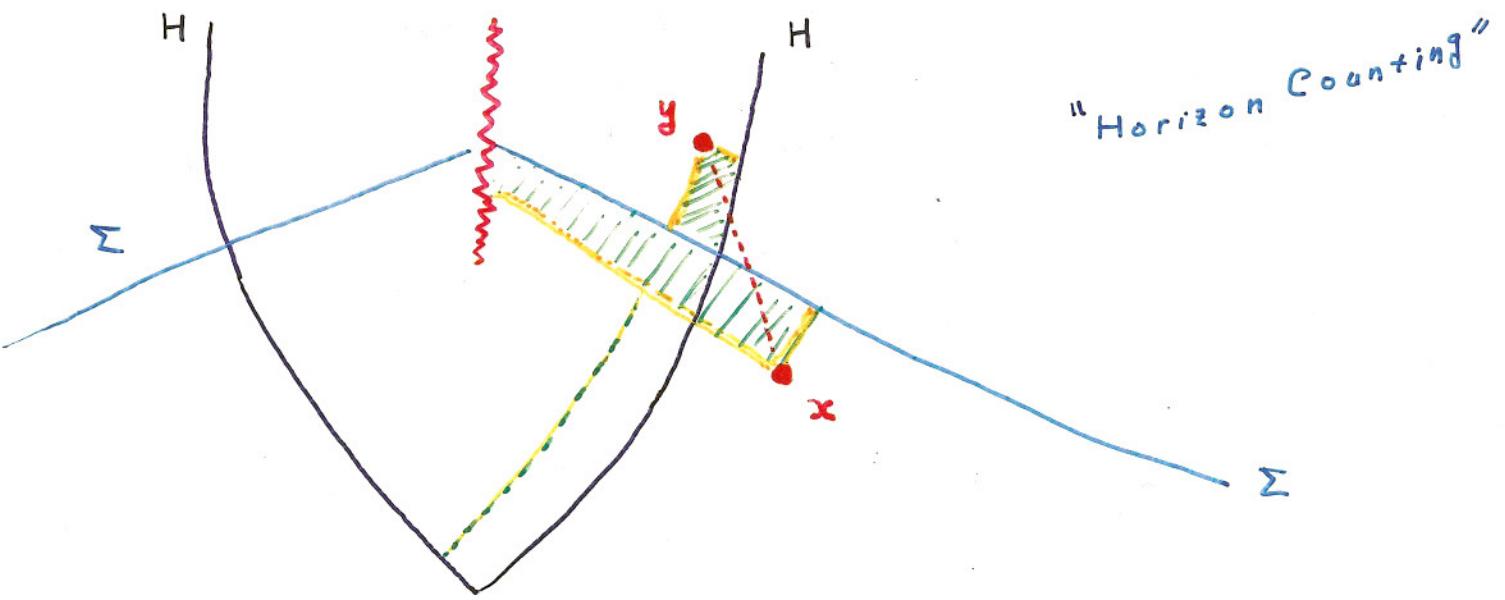
If  $\langle \Lambda \rangle = 0$  then  $\Lambda \sim \pm \frac{1}{\sqrt{V}}$

present epoch:  $V \sim 10^{240} \Rightarrow \Lambda \sim 10^{-120}$

Excludes large KK dim!

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(\*) Fourier transform argument



①  $x < \Sigma, H \quad y > \Sigma, H$

②  $x < y$  a link

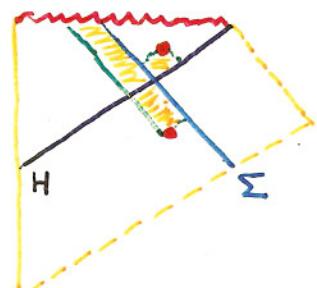
③  $x$  maximal in (past  $\Sigma$ )

$y$  minimal in (future  $\Sigma$ )  $\cap$  (future  $H$ )

③'  $x$  maximal in (past  $\Sigma$ )  $\cap$  (past  $H$ )

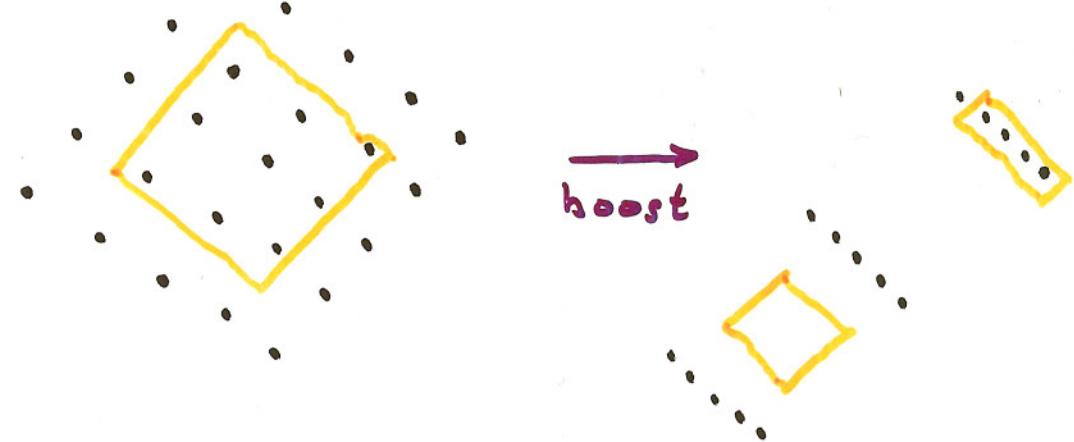
$y$  minimal in (future  $H$ )

Count such pairs  $(x, y)$



equivalent Penrose diagram  
( $\Sigma$  null)

# Why random embedding?

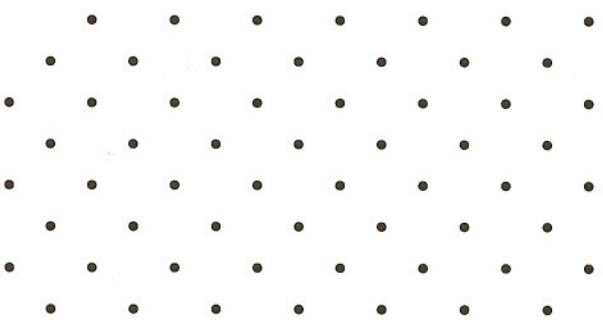


appears uniform ..... but isn't!

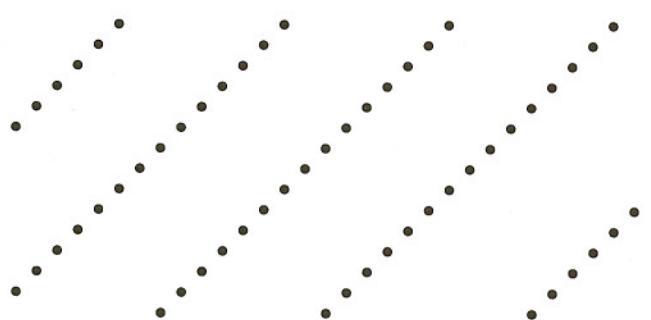
Hence the embedding is not faithful.

In contrast a Poisson process is Lorentz invariant. and uniform.

Sprinkling produces a "random lattice" and introduces a kinematic role for randomness.

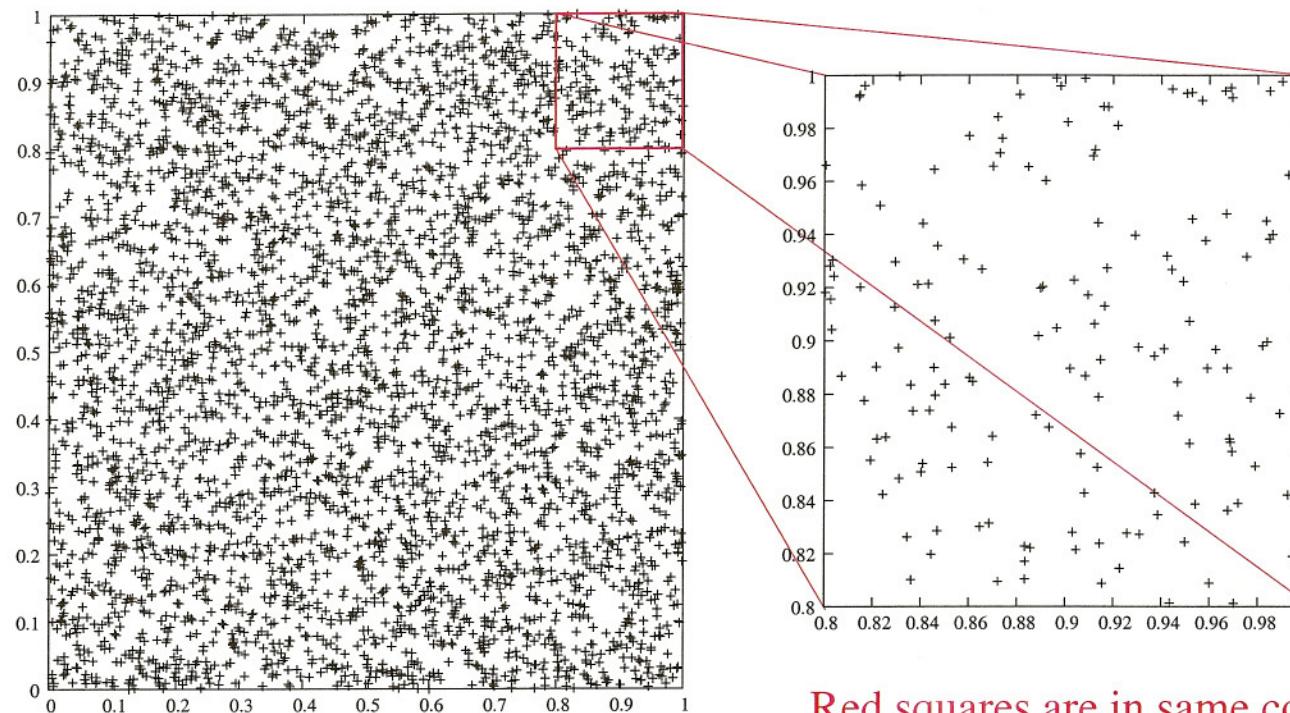


(a)

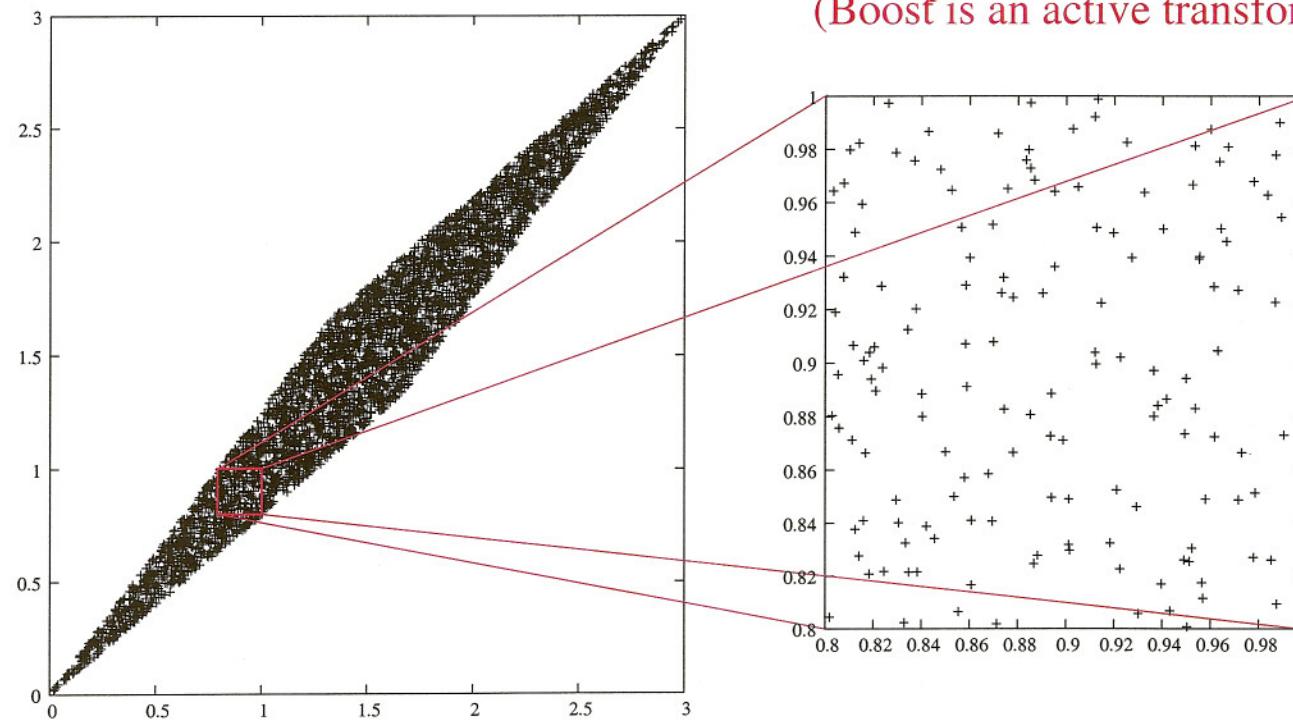


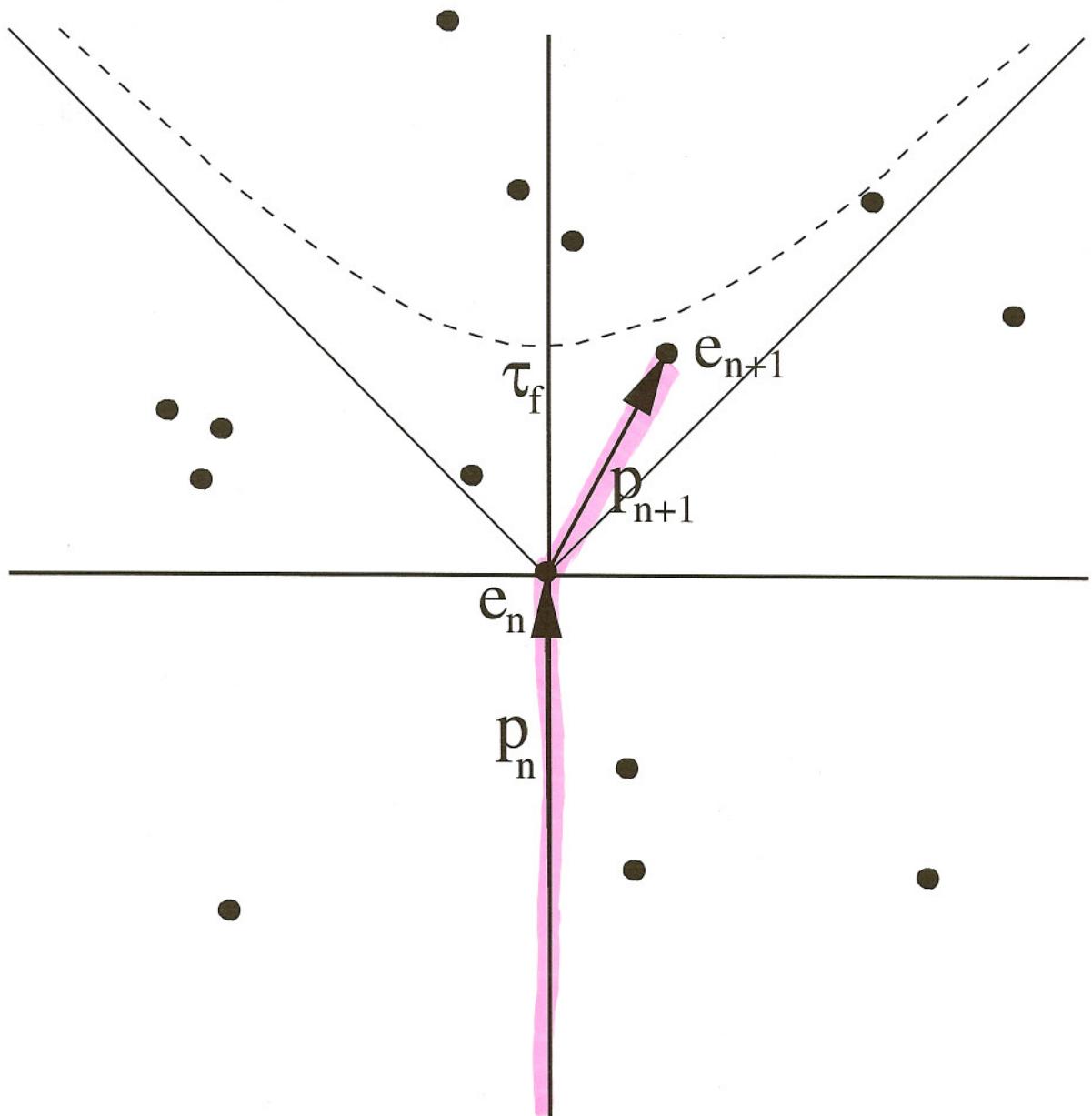
(b)

Diamond lattice before and after boost



Red squares are in same coordinate location.  
(Boost is an active transformation.)





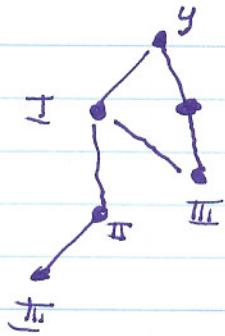
"Henson swerves" or the Lucretius effect

$$\frac{\partial \rho}{\partial \tau} = k \nabla_p^2 \rho - \frac{1}{m} p^\mu \frac{\partial}{\partial x^\mu} \rho$$

$$\frac{\partial \rho}{\partial t} = k \nabla_p^2 \left( \frac{\rho}{\sqrt{1 + p^2/m^2}} \right) - \nabla_a (w^a \rho)$$

Discrete  $\square$ -operator

$$\square \varphi(y) \propto -\frac{1}{2} \varphi(y) + \left( \sum_{\text{I}} - 2 \sum_{\text{II}} + \sum_{\text{III}} \right) \varphi(x)$$



$$\square \varphi(y) \leftrightarrow \frac{4\epsilon}{\ell^2} \left[ -\frac{1}{2} \varphi(y) + \epsilon \sum_{x \sim y} f(x) \varphi(x) \right]$$

$$f(x) = (1 - 2\tilde{\epsilon} + \frac{1}{2}\tilde{\epsilon}^2) e^{-\tilde{\epsilon}}$$

$$\tilde{\epsilon} = \epsilon | \langle x, y \rangle | \quad (\text{cardinality})$$

$$\ell^2 = \text{area / element}$$

$$(\text{nonlocality scale} = \sqrt{\frac{\ell}{\epsilon}})$$

Hilroy

Many Insights (gained)

Much Work (remains)